The impact of the spectral tail on the evolution of the kurtosis of random seas

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We perform simulations of random seas based on narrowbanded spectra with directional spreading. Our wavefields 2 are spatially homogeneous and non-stationary in time. We 3 truncate the spectral tail for the initial conditions at different 4 cut-off wavenumbers to assess the impact of the spectral tail 5 on the kurtosis and spectral evolution. We consider two cases 6 based on truncation of the wavenumber tail at $|\mathbf{k}|/k_p = 2.4$ 7 and $|\mathbf{k}|/k_p = 6$. Our simulations indicate that the peak kur-8 tosis value increases if the tail is truncated at $|\mathbf{k}|/k_p = 2.4$ 9 rather than $|\mathbf{k}|/k_p = 6$. For the case with a wavenumber cut-10 off at $|\mathbf{k}|/k_p = 2.4$, augmented kurtosis is accompanied by 11 comparatively more aggressive spectral changes including 12 redevelopment of the spectral tail. Similar trends are ob-13 served for the case with a wavenumber cut-off at $|\mathbf{k}|/k_p = 6$, 14 but the spectral changes are less substantial. Thus, the spec-15 tral tail appears to play an important role in a form of spectral 16 equilibrium that reduces spectral changes and decreases the 17 peak kurtosis value. Our findings suggest that care should 18 be taken when truncating the spectral tail for the purpose of 19 simulations/experiments. We also find that the equation of 20 Fedele (2015, J. Fluid Mech., vol. 782, pp. 25-36) provides 21 an excellent estimate of the peak kurtosis value. However, 22 the bandwidth parameter must account for the spectral tail to 23 provide accurate estimates of the peak kurtosis. 24

25 INTRODUCTION

Rogue wave occurrence in random seas and the evolution of free-surface kurtosis remain active areas of research.
Dispersive focusing based on wave components with different frequencies and directions can result in the formation of extreme waves (see, for example, Fedele et al. [1]). Nonlinear interactions between wave components can also alter the dispersive characteristics of a wave field, allowing for self

focusing (Janssen [2]). The relative importance of nonlinear interactions in the formation of rogue waves has been a focus of previous studies with comprehensive reviews [3, 4, 5, 6].

In the context of random seas, a deviation from Gaus-4 sian statistics indicates the presence of nonlinear interac-5 tions. The kurtosis of the free surface, Kur= $\langle \eta^4 \rangle / \langle \eta^2 \rangle^2$, 6 has received particular attention, as an indicator of nonlin-7 ear interactions and rogue wave occurrence (see, e.g., Mori 8 & Janssen [7]). Here, η denotes the free-surface elevation 9 and the angled brackets denote a statistical average. The ex-10 cess kurtosis, denoted as C_4 , quantifies the deviation from 11 Gaussian statistics: 12

$$C_4 = \frac{\langle \eta^4 \rangle}{3 \langle \eta^2 \rangle^2} - 1, \tag{1}$$

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yielding $C_4 = 0$ for a Gaussian process, associated with linear 13 seas. The excess kurtosis C_4 is comprised of dynamic (C_4^d) 14 and bound (C_4^b) contributions such that $C_4 = C_4^d + C_4^b$ where 15 the dynamic contribution accounts for the build-up of phase 16 correlation and the bound contribution accounts for the pres-17 ence of bound harmonics (see, e.g., [8, 9] for a more detailed 18 discussion on bound harmonics). Dynamics excess kurtosis 19 values of $C_4^d > 0$ and $C_4^d < 0$ are respectively indicative of 20 focusing and defocusing due to nonlinear interactions. 21

An analytical solution for dynamic kurtosis (C_4^d) has 22 been presented by Fedele [10], based on narrow-band di-23 rectional waves with a Gaussian-type spectrum. The ini-24 tial condition is based upon Gaussian statistics, $C_{4}^{d}(t_{0}) = 0$, 25 with random component phases and amplitudes. Fedele [10] 26 assumes that the wave field is spatially homogeneous and 27 non-stationary in time. Analysis of this problem originates 28 from Janssen [2], providing an expression for the dynamic 29 excess kurtosis of weakly nonlinear unidirectional seas. [7] 30

- extended the work of [2] based on the assumption of narrow-
- bandedness. [11] and [12] considered the role of directional 2
- effects. Fedele [10] provides an expression for dynamic kur-3
- tosis in the directional case, based on:

$$\frac{\mathrm{d}C_4^d(\tau)}{\mathrm{d}\tau} = \mathrm{BFI}^2 \frac{\mathrm{d}J}{\mathrm{d}\tau}.$$
 (2)

- Here, τ represents nondimensional time, $\tau = v^2 \omega_0 t$, where 5
- v is the spectral width and $\omega_0 = 2\pi/T_0$ is the characteristic 6
- frequency based on the characteristic wave period T_0 . The 7
- Benjamin-Feir index (BFI) is given by: 8

$$BFI = \frac{\mu\sqrt{2}}{\nu},$$
 (3)

- based on the wave steepness $\mu = k_0 \sigma$ where k_0 is the char-9 acteristic wavenumber and σ is the standard deviation of the 10 free surface, $\sigma^2 = \langle \eta^2 \rangle$. Note that the definition in (3) is a 11
- factor of $\sqrt{2}$ smaller than the one used in some other studies, 12
- e.g., Onorato et al. [13]. The function $J(\tau, R)$ in (2) depends 13
- upon the short-crestedness parameter R: 14

$$R = \frac{1}{2} \frac{\sigma_{\theta}^2}{v^2}.$$
 (4)

Here, σ_{θ} is the angular width of the spectrum which quanti-15

- fies the directional spreading of the waves. Fedele [10] cal-16
- culates the angular width σ_{θ} based on the spreading function 17

of the spectrum, $D(\theta)$: 18

$$\sigma_{\theta} = \sqrt{\frac{\int_{0}^{\pi/2} \theta^{2} D(\theta) d\theta}{\int_{0}^{\pi/2} D(\theta) d\theta}}.$$
 (5)

Using the short-crestedness parameter R and nondimensional 19 time τ , Fedele [10] found the expression: 20

$$\frac{\mathrm{d}J}{\mathrm{d}\tau} = 2\mathrm{Im}\left(\frac{1}{\sqrt{1 - 2i\tau + 3\tau^2}\sqrt{1 + 2iR\tau + 3R^2\tau^2}}\right),\qquad(6)$$

required to evaluate (2). Here, Im(x) denotes the imaginary 21 part of x. Figure 1 shows the evolution in kurtosis predicted 22 for a range of R values. As can be seen, the peak kurtosis 23 value is significantly impacted by the value of R, suggesting 24 a strong dependency on the bandwidth and spreading of the 25 waves, as well as the steepness. 26

Previous studies have investigated the evolution of kur-27 tosis for random seas, including the experiments of Ono-28 rato et al. [13] as well as the Higher-Order Spectral (HOS) 29 and Modified Nonlinear Schrödinger (MNLS) simulations of 30 Toffoli et al. [14] and Xiao et al. [15]. In this study, we per-31 form random seas simulations using the MNLS equation of 32 Trulsen et al. [16] based on an exact linear dispersion oper-33 ator. Our simulations are based on the sea-state parameters 34



Fig. 1. Dynamic excess kurtosis normalized by the square of the Benjamin-Feir index $(C_d^4/{
m BFI}^2)$ as a function of nondimensional time $(\tau = v^2 \omega_0 t)$ for different values of *R*, based on Fedele [10].

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used in Experiment B of Onorato et al. [13], a case also considered by Toffoli et al. [14] and Xiao et al. [15]. Thus, we compare our results with those of previous studies. We note that the experiments of Onorato et al. [13] are based upon wave fields which are inhomogeneous in space but stationary in time for a given location. In contrast, our simulations are based upon a spatially homogeneous random sea which is non-stationary in time, as simulated by Xiao et al. [15] and Toffoli et al. [14]. However, for narrow-banded seas with low directional spreading, mapping between space and 10 time can be performed with the group velocity, although we 11 note this is an extra approximation. Toffoli et al. [14] per-12 formed MNLS simulations of both types (spatially homoge-13 neous and non-stationary as well as spatially inhomogenous 14 and stationary) and found that the kurtosis curves agreed well 15 for the narrow-banded case considered in this study. Thus, 16 we use the experimental results of Onorato et al. [13] in our 17 comparisons. Our investigation focuses on the impact of the 18 spectral tail on kurtosis evolution, exploring the role of the 19 tail in establishing a form spectral equilibrium that reduces 20 the peak kurtosis. 21

The concept of spectral equilibrium has been considered 22 by previous studies. Non-equilibrium sea states are charac-23 terised by comparatively rapid spectral changes that eventu-24 ally slow down as the sea state moves towards a better rep-25 resentation of equilibrium for the given conditions. As dis-26 cussed by [18], wave-current interactions, sudden changes 27 in bathymetry and meteorological conditions are all possi-28 ble causes of non-equilibrium, provoking the occurrence of 29 rapid spectral changes (see, e.g., [19, 20, 21]). The investi-30 gation of Barratt et al. [22] showed that steep wave groups 31 formed under non-equilibrium conditions may exhibit aug-32 mented kinematics and a prolonged lifespan-the presence 33 of a fully-developed spectral tail was found to reduce the 34 nonlinear features of the wave groups. Physical mechanisms 35 which may impact the development of the spectral tail have 36 also been identified by previous studies. Background cur-37 rents have been shown to possibly suppress the development 38 of the spectral tail [23]. Ice sheets also tend to dissipate the 39 1 energy associated with high-wavenumber components in the

spectral tail [24, 25]. Simulations based on initial spectra that
do not include a fully-developed tail also tend to exhibit rapid
spectral evolution in the early stages (see, e.g., [26, 15]). Our

5 simulations are focused on the impact of the spectral tail on

the evolution of random seas initialised with Gaussian statis-tics.

⁸ We perform simulations based on JONSWAP spectra ⁹ truncated at different wavenumbers to alter the bandwidth ¹⁰ and prominence of the spectral tail. We monitor the conse-¹¹ quent kurtosis evolution and explain the trends based on the ¹² spectral evolution we observe. Lastly, we calculate approx-¹³ imate *R* values and compare our results to (2) to assess the ¹⁴ extent of the agreement.

15 NUMERICAL DETAILS

We perform random-sea MNLS simulations based on 16 Rayleigh distributed component amplitudes with a uniform 17 phase distribution. We consider two distinct cases, each with 18 a different cut-off wavenumber for the spectral tail. Each 19 case has been simulated a total of 20 times with a new ran-20 dom seed generated for each instance. Our analysis of the 21 spectral evolution is based upon ensemble averaging of the 22 resultant spectra. 23

24 Initial Conditions

²⁵ We define the variance density spectrum $F(\omega, \theta)$ as the ²⁶ product of a frequency spectrum $S(\omega)$ and a spreading func-²⁷ tion $D(\theta)$, where ω represents the angular frequency and θ ²⁸ represents the direction of the wave component:

$$F(\boldsymbol{\omega}, \boldsymbol{\theta}) = S(\boldsymbol{\omega})D(\boldsymbol{\theta}). \tag{7}$$

Following Onorato et al. [13], we use the JONSWAP formu lation as the frequency spectrum:

$$S(\omega) = \frac{\alpha g^2}{\omega^5} \exp\left[-\frac{5}{4} \left(\frac{\omega}{\omega_p}\right)^{-4}\right] \gamma^{\exp\left[-(\omega-\omega_p)^2/(2\sigma^2\omega_p^2)\right]},$$
(8)

where ω is the angular frequency and ω_p the peak frequency,

³² α the Phillips parameter, γ the peak enhancement factor, and ³³ the parameter σ is frequency dependent: $\sigma = 0.07$ for $\omega \le$ ³⁴ ω_p and $\sigma = 0.09$ for $\omega > \omega_p$. We use the cosine-squared ³⁵ spreading function:

$$D(\theta) = \begin{cases} \frac{2}{\Theta} \cos^2\left(\frac{\pi\theta}{\Theta}\right) & \text{ for } |\theta| \le \Theta/2. \\ 0 & \text{ for } |\theta| > \Theta/2. \end{cases}$$
(9)

Here, θ is the wave propagation direction and Θ is the direc-

³⁷ tional spreading width of the cosine-squared function. We

note the relationship between σ_{θ} in (5) and Θ in (9), given by:

$$\sigma_{\theta} = \Theta \sqrt{\frac{\pi^2 - 6}{12\pi^2}}.$$
 (10)

Table 1. Sea-state parameters.							
γ	ω_p	k_p	Θ	H_s	ε		
6.0	$0.5257 \ {\rm s}^{-1}$	$0.02796 \ {\rm m}^{-1}$	12°	11.2 m	0.16		

The product of (8) and (9) yields the variance density spectrum in the (ω, θ) coordinate system. The corresponding wavenumber spectrum in (k_x, k_y) can be calculated with the Jacobian: $\hat{S}(k_x, ky) = (1/k)(d\omega/dk)S(\omega, \theta) = (g^2/(2\omega^3))S(\omega, \theta)$, where **k** is the wavenumber vector **k** = (k_x, k_y) and we have used the deep-water dispersion relationship.

To perform random-sea simulations, we require ⁸ Rayleigh distributed component amplitudes, a_i , with expected values, μ_i , that are consistent with the defined ¹⁰ wavenumber spectrum, $\hat{S}(\mathbf{k})$. The expected amplitude for ¹¹ component \mathbf{k}_i follows from the wavenumber spectrum: ¹²

$$\mu_i = \sqrt{2\hat{S}(\boldsymbol{k}_i)}.$$
(11)

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Thus, the scale parameter of the Rayleigh distribution is 13 given by $\sqrt{2/\pi}\mu_i$ and we generate the random amplitude, 14 a_i , for component k_i using: 15

$$a_i = \sqrt{2/\pi} \,\mu_i \sqrt{-2\ln\chi},\tag{12}$$

where χ is a uniformly distributed random variable within the range [0,1] and ln is the natural logarithm. A random phase offset φ_i is also generated for each wavenumber component k_i in the range [0, 2π]. We compute the linear surface elevation at each point in space $\mathbf{x} = (x, y)$ as a superposition of the components:

$$\eta_L(\mathbf{x},t) = \sum_i a_i \cos{(\mathbf{k}_i \cdot \mathbf{x} - \omega_i t + \varphi_i)}, \quad (13)$$

using the deep-water linear dispersion relationship $\omega_i = \frac{22}{\sqrt{g|\mathbf{k}_i|}}$ to calculate the component frequencies. For the MNLS simulations, we calculate the initial complex envelope $B(\mathbf{x}, t_0)$ using the linear surface elevation η_L and the corresponding Hilbert transform η_L^H [27]: 26

$$B(\boldsymbol{x},t_0) = \left\{ \eta_L + i \eta_L^H \right\} \exp\left(-i [\boldsymbol{k}_0 \cdot \boldsymbol{x} - \boldsymbol{\omega}_0 t_0]\right).$$
(14)

Here, \mathbf{k}_0 and ω_0 represent the characteristic wavenumber and $_{27}$ frequency of the carrier wave. $_{28}$

The parameters used in this study are listed in Table 1. ²⁹ We use a peak enhancement factor (γ) of 6.0. The spectral ³⁰ peak of the JONSWAP, in terms of angular frequency (ω_p) ³¹ and wavenumber (k_p) are both listed in Table 1 (note that ω_p ³² and k_p are not simply related by the linear dispersion relation ³³ due to the presence of a Jacobian). The characteristic time ³⁴ and length scales associated with k_p are also listed in Table ³⁵

Table 2. Low-pass filter parameters for spectral tail truncation.

Case	β_1	β_2	Cut-off wavenumber
ST	2.4	20	$ \boldsymbol{k} /k_p = 2.4$
LT	6	35	$ \mathbf{k} /k_p = 6$

1 3. We use a directional spreading width (Θ) of 12°, based on

² the spreading function defined in (9), the same value as [15].

³ Our significant wave height (H_s) of 11.2 m corresponds to ⁴ a wave steepness ($\varepsilon = k_p H_s/2$) of 0.16. As calculated by

⁴ a wave steepness ($\varepsilon = k_p H_s/2$) of 0.16. As calculated by ⁵ Fedele [10], the parameters listed in Table 1 correspond to a

⁶ Benjamin-Feir index (BFI = $\mu\sqrt{2}/\nu$) of 0.78, where $\mu = \epsilon/2$

7 and v is a measure of spectral bandwidth.

We use an exponential low-pass filter to truncate the tail
 of the spectrum following [15]:

$$\Omega(|\boldsymbol{k}|/k_p, \beta_1, \beta_2) = \exp\left(-\left[\frac{|\boldsymbol{k}|}{\beta_1 k_p}\right]^{\beta_2}\right).$$
(15)

We consider two test cases labelled Case ST and Case LT, 10 where "ST" refers to a short tail and "LT" refers to a long 11 tail for the spectrum. The β_1 and β_2 values are listed in Ta-12 ble 2 together with the corresponding cut-off wavenumbers. 13 Case ST and LT feature truncation of the spectral tail at ap-14 proximately $|\mathbf{k}|/k_p = 2.4$ and $|\mathbf{k}|/k_p = 6$, respectively, based 15 on the β_1 and β_2 listed in Table 2. The resultant initial con-16 ditions are shown in Fig. 3(a) for Case ST and Fig. 4(a) for 17 Case LT. We note that approximately 21% of the total energy 18 for Case LT is associated with wavenumber components with 19 $|\mathbf{k}|/k_p > 2.4.$ 20

21 MNLS Simulations

We perform our random-sea simulations using the MNLS equation of Trulsen et al. [16], a modified version of the Trulsen & Dysthe [17] equation:

$$\frac{\partial B}{\partial t} + \mathfrak{L}B + \frac{1}{2}i\omega_0k_0^2|B|^2B + \frac{3}{2}\omega_0k_0|B|^2\frac{\partial B}{\partial x} + \frac{1}{4}\omega_0k_0B^2\frac{\partial B^*}{\partial x} + ik_0\frac{\partial\overline{\phi}}{\partial x}B = 0.$$
(16)

Here, B^* denotes the conjugate of the complex envelope 25 and $\overline{\phi}$ denotes the mean flow potential. The carrier wave 26 is aligned with the x-axis, $\mathbf{k}_0 = (k_0, 0)$, so that k_0 in (16) 27 represents the carrier wavenumber and the characteristic 28 frequency ω_0 is related to the carrier wavenumber k_0 by 29 the deep-water linear dispersion relationship, $\omega_0 = \sqrt{gk_0}$. 30 The dispersion operator \mathfrak{L} in (16) is based upon a pseudo-31 differential operator that preserves the exact linear dispersion 32

relationship, as explained by Trulsen et al. [16]:

$$\mathfrak{L}B = \frac{1}{4\pi^2} \int_{-\infty}^{\infty} i[\omega(\mathbf{k}_0 + \boldsymbol{\mu}) - \omega_0] \exp(i\boldsymbol{\mu} \cdot (\mathbf{x} - \mathbf{y})) B(\mathbf{y}, t) d\mathbf{y} d\boldsymbol{\mu}.$$
(17)

Here, $\boldsymbol{\mu} = (\lambda, \mu)$ is the modulation wavenumber. Direct nu-2 merical evaluation of (17) avoids expansion and truncation of 3 the linear dispersion relation, increasing the bandwidth limits 4 of the MNLS equation and improving the resolution of four-5 wave interactions while reducing energy leakage (see [28] 6 and [29] for a discussion on MNLS energy leakage), with 7 almost no additional computational cost. Barratt et al. [30], 8 building on [31, 32], performed a detailed comparison of the 9 exact and truncated versions of the dispersion operator for 10 focused wave groups. The MNLS equation in (17) is subject 11 to free surface and bottom boundary conditions, as well as 12 continuity for the mean flow potential $\overline{\phi}$: 13

$$\frac{\partial \overline{\phi}}{\partial z} = \frac{\omega_0}{2} \frac{\partial}{\partial x} |B|^2 \quad \text{at} \quad z = 0,$$
(18)

$$\frac{\partial \overline{\phi}}{\partial z} = 0$$
 at $z = -\infty$, (19) ¹⁴

$$\nabla^2 \overline{\phi} = 0 \quad \text{for} \quad -\infty < z < 0. \tag{20}$$

We incorporate the boundary conditions, (18) and (19), directly into the MNLS equation, (16), using the continuity condition for the mean flow, (20), as done with the fourthorder envelope equation of Janssen [33]. A single governing equation is, thus, obtained:

$$\frac{\partial B}{\partial t} + \mathfrak{L}B + \frac{1}{2}i\omega_0k_0^2|B|^2B + \frac{3}{2}\omega_0k_0|B|^2\frac{\partial B}{\partial x} + \frac{1}{4}\omega_0k_0B^2\frac{\partial B^*}{\partial x} + ik_0B\mathcal{F}^{-1}\left\{\frac{ik_x}{|\mathbf{k}|}\mathcal{F}\left\{\frac{\omega_0}{2}\frac{\partial}{\partial x}|B|^2\right\}\right\} = 0.$$
(21)

where \mathcal{F} denotes a 2D Fourier transform in x and y and 21 \mathcal{F}^{-1} denotes the inverse operation. The expression in (21) 22 is based upon the evaluation of the bottom boundary condi-23 tion (19) at $z = -\infty$ and is, therefore, a deep-water equation. 24 Thus, we obtain the initial complex envelope using (14) and 25 the envelope is marched forward in time with (21). We dis-26 cretize and numerically solve (21) using a split-step algo-27 rithm. We use spectral methods to evaluate the linear dis-28 persion operator $\mathcal{L}B$ in (17) and we use fourth-order finite 29 differencing with symmetric stencils for the spatial deriva-30 tives in the nonlinear terms. Time marching is performed 31 with the classic fourth-order Runge-Kutta scheme. The de-32 tails of the discretisation are listed in Table 3, including the 33 length (L) and width (W) of the domain. The number of grid 34

Table 3. Discre	tization Parameters.
Characteristic scales	Wavelength (λ_0) 225 m
	Wave period (T_0) 12.0s
Numerical domain	Length (<i>L</i>) 30.72 km
	Width (<i>W</i>) 20.48km
Discretisation	$N_x = 2049, \Delta x = 15 \text{ m}$
	$N_y = 1025, \Delta y = 20 \text{ m}$
	$N_t = 4501, \Delta t = 0.4 \text{ s}$

¹ points in the x-direction and y-direction are listed, denoted as

² N_x and N_y respectively, together with the corresponding grid

spacings, Δx and Δy . The size of the domain ensures 136

⁴ characteristic wavelengths (λ_0) in the *x*-direction and $91\lambda_0$ in

the *y*-direction, where $\lambda_0 = 2\pi/k_0$. The characteristic length scales of the wave envelope can be approximated with:

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$$\Lambda_x = \frac{2\pi}{k_w}, \qquad \Lambda_y = \frac{2\pi}{k_0 \sigma_{\theta}}, \qquad (22)$$

7 based on the characteristic length scales for the wavenumber

⁸ (k_0), bandwidth (k_w) and spreading parameter (ζ_0). Dimen-

9 sionless metrics for grid resolution, in the *x* and *y*-directions

10 can, thus, be defined as:

$$n_x = \frac{\Lambda_x}{\Delta x}, \qquad n_y = \frac{\Lambda_y}{\Delta y},$$
 (23)

which approximately represent the number of grid points 11 spanning the length scale of the wave envelope in the x and 12 y-directions. Based on the initial conditions, we use the peak 13 of the wavenumber spectrum, $k_0 = 0.02796 \text{ m}^{-1}$, and we use 14 (5) to obtain $\sigma_{\theta} = 0.04$ rad. We estimate the bandwidth k_w 15 for Case ST and Case LT using the spectral half-width and 16 obtain $k_w = 0.004 \text{ m}^{-1}$. Combined with the grid resolution 17 listed in Table 3, we obtain $n_x \approx 105$ and $n_y \approx 281$. The 18 simulations are time marched for a total of 150 wave peri-19 ods (T_0), where $T_0 = 2\pi/\omega_0$, with a time step (Δt) of 0.4 s. 20 Using the group velocity of the wave envelope as the char-21 acteristic velocity, we calculate a Courant-Friedrichs-Lewy 22 (CFL) condition of 0.25 for our MNLS simulations, based on 23 the discretisation parameters listed in Table 3. To assess the 24 diffusivity of our simulations, we have considered the con-25 served quantity I_2 [34]: 26

$$I_2 = \sum_{i,j} |B(x_i, y_j)|^2,$$
(24)

typically associated with energy conservation. We found the quantity I_2 to be conserved within 1% of the initial value over the entire duration of all our simulations, indicating permissibly low levels of diffusivity. We find that the $|B|^2 \partial B/\partial x$ term in (16) is particularly prone to causing simulation divergence. Thus, we apply spectral filtering to eliminate highfrequency contributions from this term—we set all compo-

nents above $|\mathbf{k}|/k_p = 5$ to zero when calculating $|B|^2 \partial B/\partial x$.

Spectral parameters

We analyse the spectral evolution of Case ST and Case 2 LT using statistical parameters to characterise the spectral 3 peak, bandwidth and directional spreading. Our selection 4 of the spectral parameters is largely based on the review 5 by Serio et al. [35]. For each simulation, we perform 6 a two-dimensional discrete Fourier transform (in x and y) 7 on the surface elevation once per wave period, and use the 8 result to calculate the variance density spectrum in terms 9 of wavenumber $S(k_x, k_y, t)$ based on a Cartesian co-ordinate 10 system: 11

$$S(k_x, k_y, t) = \frac{1}{2} |\hat{\eta}(k_x, k_y, t)|^2, \qquad (25)$$

where $\hat{\eta}$ represents the Fourier components of the surface 12 elevation. Arithmetic averaging over the ensemble ($N_i = 20$) 13 at time *t* yields the ensemble-averaged spectrum $\overline{S}(k_x, k_y, t)$: 14

$$\overline{S}(k_x, k_y, t) = \frac{1}{N_i} \sum_{i}^{N_i} S_i(k_x, k_y, t).$$
(26)

Converting to a polar co-ordinate system with the use of a Jacobian, $\overline{S}(k, \theta, t) = k\overline{S}(k_x, k_y, t)$, we characterise the directional spreading of the ensemble-averaged variance density spectrum: 18

$$\varsigma(t) = \sqrt{\frac{\sum_{j} \theta_{j}^{2} \,\overline{S}(k_{j}, \theta_{j}, t)}{\sum_{j} \overline{S}(k_{j}, \theta_{j}, t)}}.$$
(27)

Here, *k* represents the magnitude of the component wavenumber $|\mathbf{k}|$ for convenience of notation. To characterise the spectral peak and the bandwidth, we calculate the frequency spectrum $\overline{S}(f,\theta,t) = J \overline{S}(k,\theta,t)$ where $J = 4\pi\sqrt{k/g}$. Integration over θ yields the omnidirectional frequency spectrum $\overline{S}(f,t)$ used to estimate the peak frequency:

$$f_p(t) = \frac{\sum_j f_j [\overline{S}(f_j, t)]^4}{\sum_j [\overline{S}(f_j, t)]^4},$$
(28)

based on the omnidirectional frequency spectrum raised to the fourth power, as recommended by Young [36]. We also estimate the bandwidth based on the omnidirectional frequency spectrum, using the peakedness parameter introduced by Goda [37]:

$$Q_p(t) = \frac{2}{m_0^2} \int_0^\infty f[\overline{S}(f,t)]^2 \mathrm{d}f, \qquad (29)$$

where,

$$m_0 = \int_0^\infty \overline{S}(f, t_0) \mathrm{d}f. \tag{30}$$



Fig. 2. Kurtosis evolution for Case ST (red line) and Case LT (blue line) compared against other studies, including (\triangle) the experiments of Onorato et al. [13]. The shaded grey bands represent 95% confidence intervals. The simulation results of Xiao et al. [15] are shown: (---) MNLS and (\longrightarrow) HOS as well as the simulation results of Toffoli et al. [14]: (\bigcirc) MNLS and (+) HOS. All the results are based upon JONSWAP spectra ($\gamma = 6$) with steepness $\varepsilon = 0.16$ and a Benjamin-Feir index (BFI) of 0.78 based on the definition in (3).

We use the trapezoidal method to perform the numerical inte-

² gration in (29) based on unequal point spacing. Our estimate

 $_3$ of spectral bandwidth (v) relates inversely to the peakedness

⁴ parameter Q_p :

$$\mathbf{v}(t) = \frac{1}{\sqrt{\pi}Q_p},\tag{31}$$

consistent with the bandwidth metric used by Serio et al. [35]
to calculate the Benjamin-Feir index (BFI). We also use v in
(31) as our bandwidth metric when calculating the BFI. The
spectral parameters defined in (27), (28) and (31) thus form
the basis of our spectral evolution analysis.

10 RESULTS AND DISCUSSION

We analyse the kurtosis evolution for Case ST and Case LT and explain the observations based on the spectral evolution, using contour plots of the ensemble-averaged spectra as well as the parameters defined in (27), (28) and (31). Lastly, we compare the simulation results for kurtosis with the theory of Fedele [10] and we briefly discuss the selection of an appropriate bandwidth parameter to characterise each case.

18 Kurtosis Evolution

The evolution of kurtosis for our MNLS simulations is 19 shown in Fig. 2, including both dynamic and bound contribu-20 tions up to the third order. The shaded grey bands represent 21 95% confidence intervals for the ensemble-averaged MNLS 22 results. The experimental results of Onorato et al. [13] are 23 also shown. Both Toffoli et al. [14] and Xiao et al. [15] per-24 formed similar simulations to those in this study, using the 25 MNLS equation as well as a Higher-Order Spectral (HOS) 26

code, and the results are depicted in Fig. 2. As discussed in the introduction, the results of Onorato et al. [13] are based 2 on waves propagating along a tank. We perform space/time 3 mapping with the group velocity for the purposes of comparing our simulation results to the experiments of Onorato et 5 al. [13]. The x-axis in Fig. 2 shows the corresponding spa-6 tial x/λ_0 or temporal $t/(2T_0)$ parameter with kurtosis shown 7 on the y-axis (excluding the contribution of bound harmon-8 ics). Here, λ_0 and T_0 represent the characteristic wavelength 9 and wave period, respectively. We see good agreement be-10 tween the MNLS simulations results of Toffoli et al. [14], 11 Xiao et al. [15] and Case ST of this study. A peak kurto-12 sis value of 3.89 is observed for Case ST and agreement be-13 tween the MNLS simulations appears to be particularly good 14 in the vicinity of the peak. Similar to Case ST of this study, 15 the MNLS simulations of both Toffoli et al. [14] and Xiao 16 et al. [15] effectively truncated the wavenumber spectrum 17 of the surface elevation at $|\mathbf{k}|/k_p = 2$, by limiting the mod-18 ulation wavenumber of the envelope to $|\boldsymbol{\mu}|/k_p \leq 1$. Here, 19 $\boldsymbol{\mu} = (\lambda, \mu)$ is the modulation wavenumber defined relative 20 to the wavenumber of the carrier wave, $\mathbf{k} = (k_p, 0)$. Thus, 21 Fig. 2 also serves to verify our simulations. The HOS results 22 of Toffoli et al. [14] and Xiao et al. [15] differ, however, 23 from the MNLS results for Case ST and agree better with the 24 MNLS results for Case LT as well as the experimental results 25 of Onorato et al. [13]. Case LT is based upon a spectral tail 26 truncated at $|\mathbf{k}|/k_p = 6$ and, thus, features a more prominent 27 spectral tail. Likewise, the experiments of Onorato et al. [13] 28 and the HOS simulations of Toffoli et al. [14] and Xiao et al. 29 [15] all included a fully-developed spectral tail in the initial 30 conditions. Thus, the differences between Case ST and Case 31 LT appear to be the result of the spectral tail and the findings 32 are consistent other studies. Case LT, in this study, reaches a 33 peak kurtosis value of 3.52, approximately 10% lower than 34 the peak kurtosis value for Case ST. Inclusion of the spectral 35 tail up to $|\mathbf{k}|/k_p = 6$ in the initial conditions, thus, appears to 36 reduce the peak kurtosis value while artificial truncation of 37 the spectral tail at $|\mathbf{k}|/k_p = 2.4$ augments the peak kurtosis 38 value. The relatively good agreement between Case LT and 39 the experiments/HOS results also suggests that the MNLS 40 equation provides better kurtosis estimates if the spectral tail 41 is included, despite the narrow-bandwidth limitations of the 42 equation. 43

Spectral Evolution

To clarify the trends in kurtosis observed in the previ-45 ous section, we have analysed the evolution of the ensemble-46 averaged variance density spectrum, $\overline{S}(k_x, k_y)$, for Case ST 47 and Case LT with the results shown in Fig. 3 and 4 respec-48 tively. The wavenumbers k_x and k_y have been normalised by 49 the characteristic wavenumber $k_0 = 0.02796 \text{ m}^{-1}$, the initial 50 spectral peak listed in Table 1. The contour plots in each 51 figure are shown at times: (a) $t/T_0 = 0$; (b) $t/T_0 = 50$; and 52 (c) $t/T_0 = 100$. Note that the contour levels are logarithmi-53 cally distributed. As can be seen in Fig. 3(a), Case ST fea-54 tures a narrow-banded spectrum with truncation of the tail in 55 the vicinity of $|\mathbf{k}|/k_p = 2.4$. Fig. 3(b) and 3(c) reveal the 56



Fig. 3. Contour plots of the ensemble-averaged variance density spectrum $\overline{S}(k_x, k_y)$ for Case ST featuring truncation of the spectral tail in the vicinity of $|\mathbf{k}|/k_p = 2.4$: (a) $t/T_0 = 0$; (b) $t/T_0 = 50$; (c) $t/T_0 = 100$. The contour levels are logarithmic, ranging from 1×10^{-5} to 1×10^{-2} .



Fig. 4. Contour plots of the ensemble-averaged variance density spectrum $\overline{S}(k_x, k_y)$ for Case LT featuring truncation of the spectral tail in the vicinity of $|\mathbf{k}|/k_p = 6$: (a) $t/T_0 = 0$; (b) $t/T_0 = 50$; (c) $t/T_0 = 100$. The contour levels are logarithmic, ranging from 1×10^{-5} to 1×10^{-2} .



Fig. 5. Evolution of ensemble-averaged spectral parameters for Case ST (red line) and Case LT (blue line): (a) spectral bandwidth v defined in (31); (b) spreading parameter ς defined in (27); (c) peak frequency f_p defined in (28).

¹ rapid broadening of the spectrum which occurs during the ² simulation—the truncated tail partially redevelops and the ³ directional spreading increases. Similar features are appar-⁴ ent for Case LT, shown in Fig. 4. Case LT features a more ⁵ prominent spectral tail for the initial conditions since trun-⁶ cation is performed in the vicinity of $|\mathbf{k}|/k_p = 6$. However, ⁷ rapid broadening of the spectrum over time is also observed ⁸ for Case LT, although less directional spreading is apparent in Fig. 4(b) and 4(c) than the corresponding plots in Fig. 3. The differences in spectral evolution between Case ST and Case LT are best captured by the statistical parameters defined in (27), (28) and (31).

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Figure 5 shows the evolution of the ensemble-averaged spectral parameters for Case ST and Case LT. The spectral bandwidth parameter v, defined in (31), is shown in Fig. 5(a). 7 The spreading parameter ζ , defined in (27), is shown in Fig. 8



Fig. 6. Kurtosis evolution for Case ST (red line) and Case LT (blue line) compared against the solution of Fedele [10] based on different bandwidth parameters: (dashed, ---) bandwidth based on spectral half-width for both Case ST and Case LT yields R = 0.031; (dotted, ---) bandwidth based on (31) for Case ST yields R = 0.052; (solid, ---) bandwidth based on (31) for Case LT yields R = 0.022. The shaded grey bands represent 95% confidence intervals for the ensemble-averaged results.

5(b). The peak frequency f_p , defined in (28), is shown in 1 Fig. 5(c). All of the parameters have been normalised by 2 their initial value, at time t_0 . Figure 5(a) shows that Case ST 3 and Case LT both exhibit an increase in spectral bandwidth, 4 however the increase in v/v_0 is more rapid for Case ST and 5 the final value of 1.85 exceeds the final value of 1.48 for Case 6 LT by approximately 25%. Similarly, the spreading param-7 eter in Fig. 5(b) also exhibits an increase in ζ/ζ_0 for both 8 cases-the values are similar towards the start and end of the 9 simulations however Case ST exhibits a more rapid increase 10 in between. Lastly, Fig. 5(c) demonstrates a reduction in 11 the peak frequency for both cases, consistent with the down-12 shift of the spectral peak observed in other studies (see, e.g., 13 [36-37]). The frequency downshift is also observed to occur 14 more rapidly for Case ST than Case LT and the final value of 15 0.903 for Case ST is approximately 4% lower than the final 16 value of 0.935 observed for Case LT. Thus, all the spectral 17 parameters indicate that the spectral changes for Case ST oc-18 cur more rapidly and are more pronounced than those of Case 19 LT. Inclusion of a more prominent spectral tail in Case LT 20 thus appears to reduces the spectral changes observed dur-21 ing the simulations. Truncation of the spectral tail close to 22 the spectral peak in Case ST conversely augments the spec-23 tral changes which occur during the simulations. Thus, the 24 spectral tail appears to play an important role in establishing 25 the spectral equilibrium of the sea state and care should be 26 taken when truncating the tail in a simulation or laboratory 27 setting. The kurtosis results shown in Fig. 2 indicate that the 28 rapid spectral changes in Case ST augment the peak kurtosis 29 value, relative to Case LT, demonstrating the importance of 30 the tail in determining the peak kurtosis. 31

Comparison with Theory

We compare our ensemble-averaged kurtosis results 2 with the solution of Fedele [10], with the results shown in 3 Fig. 6. The kurtosis curves for Case ST and Case LT are 4 both shown, expressed as dynamic excess kurtosis, see (1), 5 normalised by the square of the Benjamin-Feir index, see 6 (3). The kurtosis curves have been plotted against non-7 dimensional time, $\tau = v^2 \omega_0 t$, based on the spectral band-8 width (v) and the characteristic frequency (ω_0). Our calcu-9 lation of the shortcrestedness parameter R is based upon (4) 10 using the angular width in (5). The dashed line in Fig. 6, 11 calculated by Fedele [10], bases the bandwidth parameter v12 on the spectral half-width and, thus, yields R = 0.032 with 13 the same curve for Case ST and Case LT (since the spec-14 tral half-width is not altered by truncation of the tail). The 15 solid (R = 0.022) and dotted (R = 0.052) lines in Fig. 6 are 16 based on the bandwidth parameter in (31), which does ac-17 count for truncation of the spectral tail. As can be seen, 18 the kurtosis curves based on (31) provide better agreement 19 with the simulation results, compared with the curve based 20 on the spectral half-width. We note that the peak kurtosis 21 value is particularly well predicted, although the long-term 22 behaviour differs-the kurtosis decline after the peak occurs 23 faster for the simulations than predicted by the theoretical 24 results. Thus, we find that Fedele [10] provides an excel-25 lent estimate for the peak kurtosis value in our simulations. 26 However, the bandwidth parameter v must account for the 27 spectral tail to accurately predict the kurtosis peak. We find 28 that the bandwidth parameter in (31), based on the peaked-29 ness parameter of Goda [37], appears to be suitable for this 30 purpose, consistent with the recommendations of Serio et al. 31 [35]. 32

CONCLUSION

Our findings indicate that artificial truncation of the 34 spectral tail augments the peak kurtosis value of a random 35 sea initialised with Gaussian statistics. Truncation of the 36 tail results in more aggressive spectral changes during the 37 simulation, characterised by spectral broadening in terms of 38 bandwidth and spreading as well as downshifting of the spec-39 tral peak. The spectral tail is also observed to redevelop 40 during the course of the simulation. Thus, the spectral tail 41 appears to play an important role in establishing a form of 42 spectral equilibrium that reduces spectral changes and de-43 creases the peak kurtosis value. Care should, thus, be taken 44 when artificially truncating the tail for the purpose of sim-45 ulations/experiments. We find that the MNLS equation of 46 Trulsen et al. [16] can be used to estimate the peak kurtosis 47 value, by including the spectral tail in the initial conditions, 48 despite the bandwidth limits of the equation. 49

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